## Current Issues in Modeling Disruption Effects on Plasma Facing Components

Ahmed Hassanein
(Presented by Rich Mattas)
Argonne National Laboratory

Presented at the APEX Study Meeting November 8-11, 1999, UCLA, Ca.

## **Current Concerns and Studies**

### Vapor-Cloud Effects

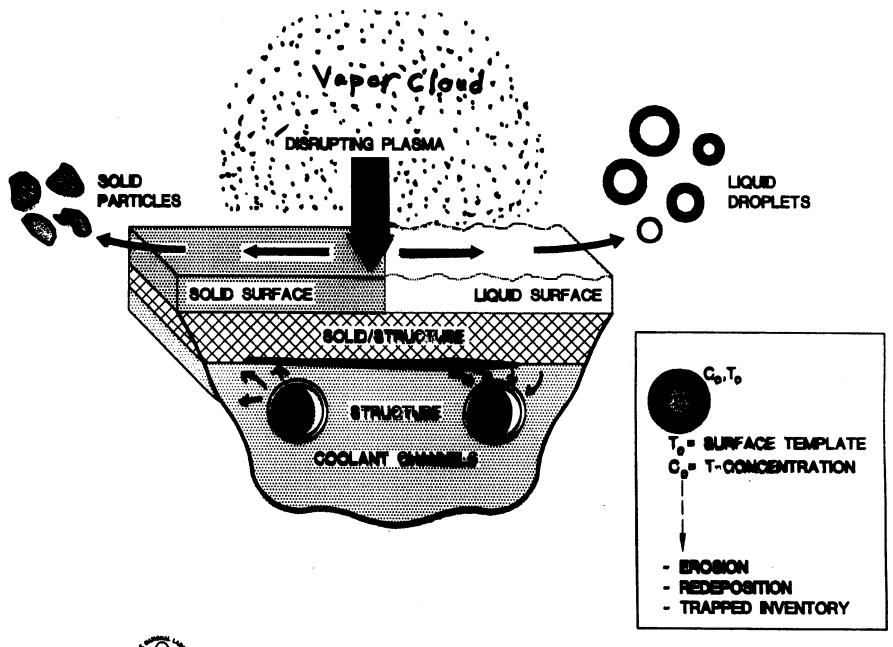
- Vapor turbulence and MHD instabilities
- Expansion and damage to nearby components:
- Impurities effect on energy radiation

### Secondary Radiation

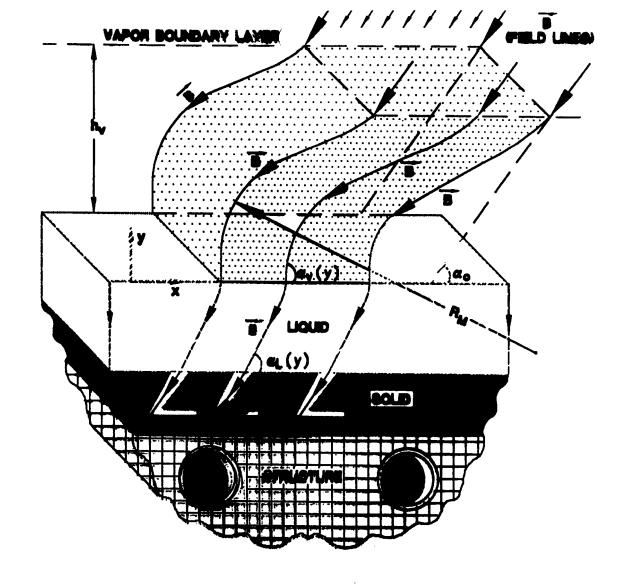
- Transport and damage to nearby components
- Less "secondary" vapor-shielding effect

### Macroscopic Erosion

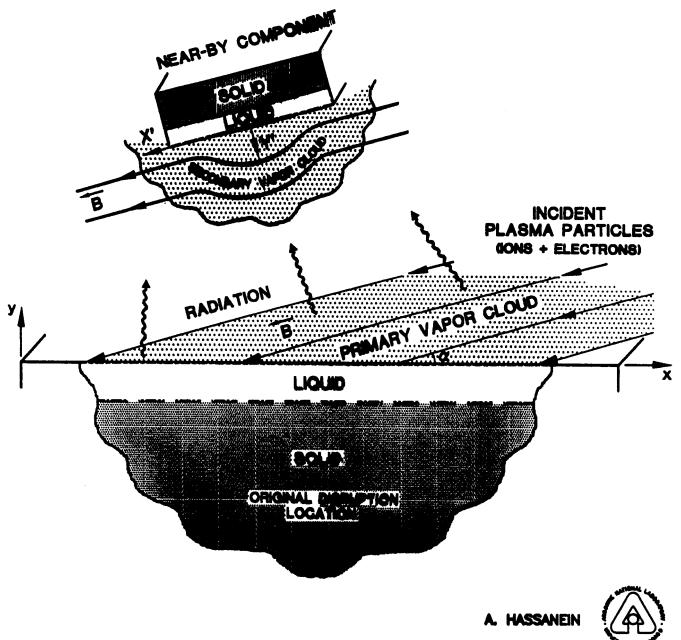
- Splashing and transport to other components
- MHD instabilities in liquid layer
- Brittle destruction of carbon-based materials



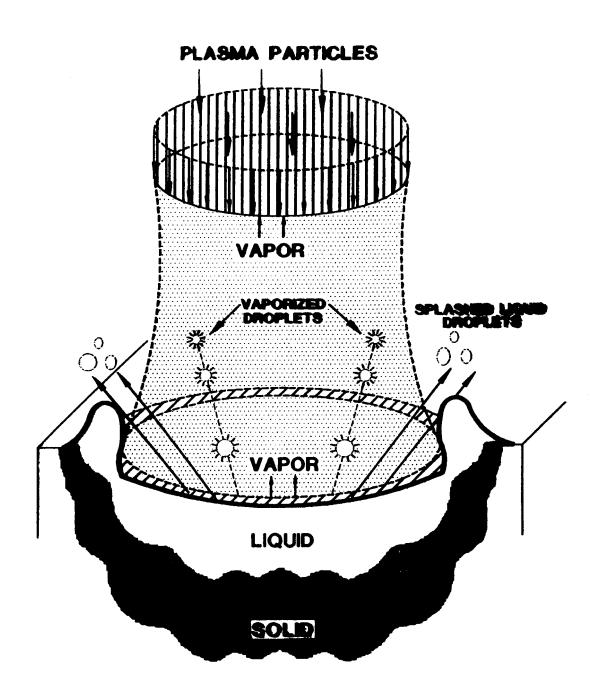
A. HASSANEIN











## **Droplet Shielding Effect**

ullet Heating of target surface by radiation power W, emits macroscopic particles with density  $n_d$ , and radius  $R_{d0}$  leaving the surface with velocity U

$$n_{do} = \frac{W_s}{q_d V_{do} U}$$
 ;  $V_{do} = \frac{4}{3} \pi R_{do}^3$ 

◆ Because of radiation absorption by macroscopic particles radiation power to surface decreases

$$\frac{dW}{dx} = -\frac{W}{l_v}$$
;  $l_v = \frac{1}{n_d \sigma}$ ;  $\sigma = \zeta \pi R_d^2$ 

♦ Absorbed radiation is spent in vaporization of macroscopic particles reducing their volume V₄ and therefore radiation power decreases

$$U\frac{dV_d}{dx} = -\frac{W}{4}\zeta R_d^2$$

• Net radiation power reaching target surface depends only on ratio  $\xi$  between vaporization energy q, and destruction energy  $q_d$ 

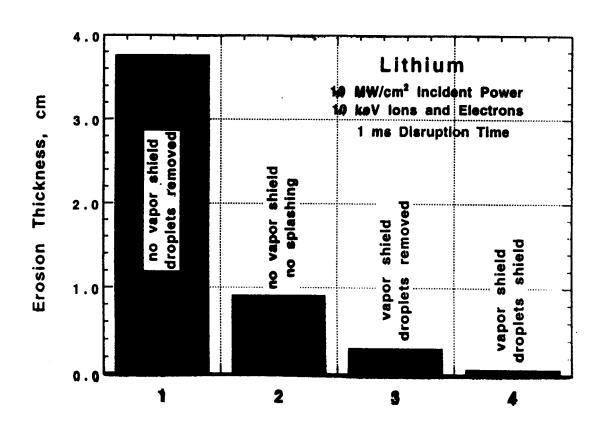
$$\xi = \frac{W_s}{W_0} = \frac{1}{1+\lambda}$$
,  $\lambda = \frac{q_v}{q_d}$ 

ullet Macroscopic particles are entirely vaporized at distance L depending on initial size of MP  $R_{do}$ , velocity U, radiation power from vapor cloud toward to the surface  $W_o$ , and ratio  $\xi$ ,

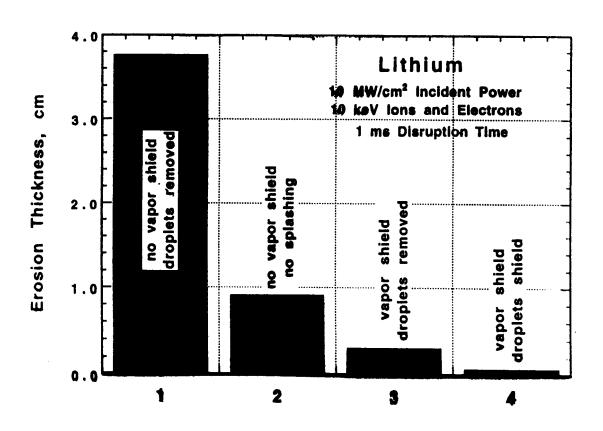
$$\frac{L}{R_{do}} = \frac{q_{\nu}U}{W_o} \frac{4}{\zeta} F , \quad F = \frac{v_0}{u_0},$$

$$v_o = \frac{1}{3} \ln \frac{\sqrt{1 + u_0 + u_0^2}}{1 - u_0} + \frac{1}{\sqrt{3}} arctg \frac{u_0 \sqrt{3}}{2 + u_0}$$
,  $u_o = \sqrt[3]{\frac{\lambda}{1 + \lambda}}$ 

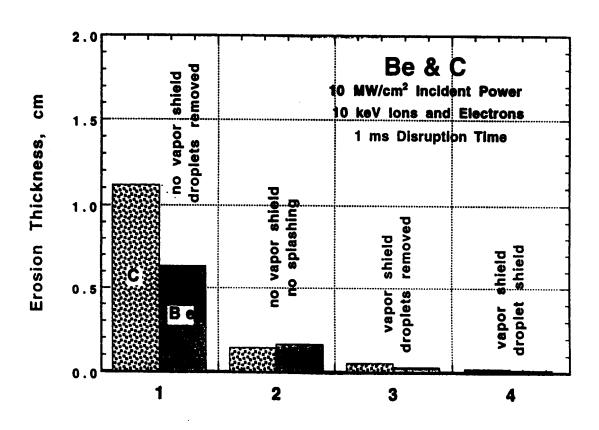
# HEIGHTS Analysis of Total Erosion Thickness of Divertor Plate Materials



# HEIGHTS Analysis of Total Erosion Thickness of Divertor Plate Materials



# HEIGHTS Analysis of Total Erosion Thickness of Divertor Plate Materials



## APEX LIQUID WALLS RESPONSE DURING TOKAMAK PLASMA DISRUPTIONS

· Main mass loss mechanisms:

**Vaporization** 

Ablation - macroscopic particles (droplets)

• Processes that decrease net mass loss of liquid walls:

#### 1) Vapor shield

- Vapor cloud absorbs incoming energy flux Wo
- Conversion to back radiation (secondary radiation)
- Decrease heat load on liquid surface to < 10% Wo

#### 2) droplet shield

- Due to ablation (splashing) the surface emits droplets therefore cloud of vapor and macroscopic particles exists nearby the exposed surface
- droplets (macroscopic particles) absorb radiation power and decrease heat load onto surface
- droplets vaporization path length depends on vapor dynamics in oblique magnetic field

### PROPOSED APEX WORK

#### A. HASSANEIN (ANL)

- (1) Disruption effects on liquid metal surfaces.
  - -- Vporization
  - -- Splashing
  - -- Effects on design and other nearby components.
  - -- Condensation on diagnostic windows (important issue)
- (2) Disruption effects on FLIBE surfaces.
  - All similar issues as above.
- (3) Disruption effects on NSTX device.
  - -- Calculations for actual design parameters of NSTX.
- (4) Particle pumping of liquid metal and FLIBE surfaces.
  - Develop kinetic models to study particle behaviour and recycling for moving liquid surfaces. (important issue)
  - -- This include both D/T and He particles (self-pumping?)
  - -- This issue significantly affect SOL and plasma interface behavior
- (5) MHD effects on free liquid-metal surfaces
  - -- We have the capabilities to study the various effects of strong B-field on free moving liquid jets and surfaces. This work will be complementary to work done at UCLA and others if needed. We are doing this for the High-Energy Physics program.